Lecture 12

e Structure Const.

Atomic structure

Atomic structure: background

 Our studies of hydrogen-like atoms revealed that the spectrum of the Hamiltonian,

$$\hat{H}_0 = \frac{\hat{\mathbf{p}}^2}{2m} - \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r}$$

is characterized by large n^2 -fold degeneracy.

- However, although the non-relativistic Schrödinger Hamiltonian provides a useful platform, the formulation is a little too naïve.
- The Hamiltonian is subject to several classes of "corrections", which lead to important physical ramifications (which reach beyond the realm of atomic physics).
- In this lecture, we outline these effects, before moving on to discuss multi-electron atoms in the next.

Atomic structure: hydrogen atom revisited

- As with any centrally symmetric potential, stationary solutions of \hat{H}_0 index by quantum numbers $n\ell m$, $\psi_{n\ell m}(\mathbf{r}) = R_{n\ell}(r)Y_{\ell m}(\theta,\phi)$.
- For atomic hydrogen, n^2 -degenerate energy levels set by

$$E_n = -\operatorname{Ry} \frac{1}{n^2}, \qquad \operatorname{Ry} = \left(\frac{e^2}{4\pi\epsilon_0}\right)^2 \frac{m}{2\hbar^2} = \frac{e^2}{4\pi\epsilon_0} \frac{1}{2a_0}$$

where *m* is reduced mass (ca. electron mass), and $a_0 = \frac{4\pi\epsilon_0}{e^2}\frac{\hbar^2}{m}$.

- For higher single-electron ions (He⁺, Li²⁺, etc.), $E_n = -Z^2 \frac{Ry}{n^2}$.
- Allowed combinations of quantum numbers:

n	ℓ	Subshell(s)
1	0	1 <i>s</i>
2	0, 1	2s2p
3	0, 1, 2	3s 3p 3d
n	$0\cdots(n-1)$	$ns\cdots$

Atomic structure: hydrogen atom revisited

• However, treatment of hydrogen atom inherently non-relativistic:

$$\hat{H}_0 = \frac{\hat{\mathbf{p}}^2}{2m} - \frac{1}{4\pi\epsilon_0}\frac{Ze^2}{r}$$

is only the leading term in relativistic treatment (Dirac theory).

- Such relativistic corrections begin to impact when the electron becomes relativistic, i.e. $v \sim c$.
- Since, for Coulomb potential, $2\langle k.e. \rangle = -\langle p.e. \rangle$ (virial theorem), $\frac{1}{2}mv^2 = \langle k.e. \rangle = -E_{000} = Z^2 Ry$. Using identity,

$$Z^2 \mathrm{Ry} = \frac{1}{2} mc^2 (Z\alpha)^2, \qquad \alpha = \frac{e^2}{4\pi\epsilon_0} \frac{1}{\hbar c} \simeq \frac{1}{137}$$

where α denotes the fine structure constant, we find

$$\frac{\mathbf{v}}{\mathbf{c}} = \mathbf{Z}\alpha$$

The "real" hydrogen atom: outline

- Terms of higher order in $\frac{v}{c} = Z\alpha$ provide relativistic corrections which lead to lifting of the degeneracy.
- These corrections (known as **fine-structure**) derive from three (superficially) different sources:
 - (a) relativistic corrections to the kinetic energy;
 - (b) coupling between spin and orbital degrees of freedom;
 - (c) and a contribution known as the Darwin term.

(a) Relativistic corrections to kinetic energy

• From the relativistic energy-momentum invariant,

$$E = \sqrt{\mathbf{p}^2 c^2 + m^2 c^4} = mc^2 + \frac{\mathbf{p}^2}{2m} - \frac{1}{8} \frac{(\mathbf{p}^2)^2}{m^3 c^2} + \cdots,$$

we can anticipate the leading correction to the non-relativistic Hamiltonian is given by

$$\hat{H}_1 = -\frac{1}{8} \frac{(\hat{\mathbf{p}}^2)^2}{m^3 c^2}$$

• The relative scale of perturbation

$$\frac{\langle \hat{H}_1 \rangle}{\langle \hat{H}_0 \rangle} \simeq \frac{p^2}{m^2 c^2} = \frac{v^2}{c^2} \simeq (Z\alpha)^2$$

where $\alpha = \frac{e^2}{4\pi\epsilon_0} \frac{1}{\hbar c} \simeq \frac{1}{137}$. i.e. \hat{H}_1 is only a small perturbation for small atomic number, Z.

(a) Relativistic corrections to kinetic energy

$$\hat{H}_1 = -\frac{1}{8} \frac{(\hat{\mathbf{p}}^2)^2}{m^3 c^2}$$

• Since
$$[\hat{H}_1, \hat{\mathbf{L}}^2] = 0$$
 and $[\hat{H}_1, \hat{\mathcal{L}}_z] = 0$,
 $\langle n\ell m | [\hat{H}_1, \hat{\mathbf{L}}^2] | n\ell' m' \rangle = \hbar^2 [\ell'(\ell'+1) - \ell(\ell+1)] \langle n\ell m | \hat{H}_1 | n\ell' m' \rangle = 0$
 $\langle n\ell m | [\hat{H}_1, \hat{\mathcal{L}}_z] | n\ell' m' \rangle = \hbar(m'-m) \langle n\ell m | \hat{H}_1 | n\ell' m' \rangle = 0$

• Therefore, the off-diagonal matrix elements vanish:

$$\langle n\ell m | \hat{H}_1 | n\ell' m' \rangle = 0 \quad \text{for} \quad \ell \neq \ell' \quad \text{or} \quad m \neq m'$$

and we can estimate energy shift without having to invoke degenerate perturbation theory.

(a) Relativistic corrections to kinetic energy

• Making use of the identity,

$$\hat{H}_1 = -\frac{1}{8} \frac{(\hat{\mathbf{p}}^2)^2}{m^3 c^2} = -\frac{1}{2mc^2} \left[\hat{H}_0 - V(r) \right]^2, \qquad V(r) = -\frac{Ze^2}{4\pi\epsilon_0} \frac{1}{r}$$

scale of resulting energy shift can be obtained from first order perturbation theory,

$$\langle n\ell m | \hat{H}_1 | n\ell m \rangle = -\frac{1}{2mc^2} \left(E_n^2 - 2E_n \langle V(r) \rangle_{n\ell} + \langle V^2(r) \rangle_{n\ell} \right)$$

• Using the identities,

$$\left\langle \frac{1}{r} \right\rangle_{n\ell} = \frac{Z}{a_0 n^2}, \qquad \left\langle \frac{1}{r^2} \right\rangle_{n\ell} = \frac{Z^2}{a_0^2 n^3 (\ell + 1/2)}$$

resulting energy shift acquires angular momentum dependence:

$$\langle \hat{H}_1 \rangle_{n\ell m} = -\frac{mc^2}{2} \left(\frac{Z\alpha}{n}\right)^4 \left(\frac{n}{\ell+1/2} - \frac{3}{4}\right)$$

- Spin degree of freedom of electron emerges naturally from relativistic formulation of quantum mechanics. Alongside the spin, this formulation leads to a further relativistic correction which involves coupling between spin and orbital degrees of freedom.
- For a general potential V(r), this spin-orbit coupling is given by:

$$\hat{H}_2 = \frac{1}{2m^2c^2} \frac{1}{r} (\partial_r V) \,\hat{\mathbf{L}} \cdot \hat{\mathbf{S}}$$

• For a hydrogen-like atom, $V(r) = -\frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r}$ and

$$\hat{H}_2 = \frac{1}{2m^2c^2} \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r^3} \hat{\mathbf{L}} \cdot \hat{\mathbf{S}}$$

(b) Spin-orbit coupling: physical origin

- Physically, as electron moves through electric field of nucleus, $\mathbf{E} = -\nabla V(r) = -\hat{\mathbf{e}}_r(\partial_r V)$, in its rest frame it will experience a magnetic field, $\mathbf{B} = \frac{1}{c^2} \mathbf{v} \times \mathbf{E}$.
- In this field, the spin magnetic moment of the electron, $\mu_s = -\frac{e}{m}\mathbf{S}$, leads to an additional interaction energy,

$$-\boldsymbol{\mu}_{s} \cdot \mathbf{B} = -\frac{e}{(mc)^{2}} \mathbf{S} \cdot (\mathbf{p} \times \hat{\mathbf{e}}_{r}(\partial_{r} V)) = \frac{e}{(mc)^{2}} \frac{1}{r} (\partial_{r} V) \mathbf{L} \cdot \mathbf{S}$$

where we have used the relation $\mathbf{p} \times \hat{\mathbf{e}}_r = -\frac{1}{r} \mathbf{L}$.

- Additional factor of 1/2 derives from further relativistic effect known as Thomas precession.
- Those discontent with heuristic derviation need only wait for Dirac formulation...

$$\hat{H}_2 = \frac{1}{2m^2c^2} \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r^3} \hat{\mathbf{L}} \cdot \hat{\mathbf{S}}$$

- Without spin-orbit interaction, eigenstates of hydrogen-like atoms can be expressed in basis of mutually commuting operators, *Ĥ*₀, *L*², *L̂_z*, *Ŝ*², and *Ŝ_z*.
- However, with spin-orbit, total Hamiltonian no longer commutes with \hat{L}_z or \hat{S}_z – useful to exploit degeneracy of \hat{H}_0 to switch to new basis in which $\hat{\mathbf{L}} \cdot \hat{\mathbf{S}}$ is diagonal.
- Achieved by turning to basis of eigenstates of the operators, \hat{H}_0 , \hat{J}^2 , \hat{J}_z , \hat{L}^2 , and \hat{S}^2 , where $\hat{J} = \hat{L} + \hat{S}$. Since $\hat{J}^2 = \hat{L}^2 + \hat{S}^2 + 2\hat{L} \cdot \hat{S}$, it follows that,

$$\hat{\mathbf{L}} \cdot \hat{\mathbf{S}} = \frac{1}{2}(\hat{\mathbf{J}}^2 - \hat{\mathbf{L}}^2 - \hat{\mathbf{S}}^2) = \frac{1}{2}(j(j+1) - \ell(\ell+1) - s(s+1))$$

$$\hat{\mathbf{L}} \cdot \hat{\mathbf{S}} = \frac{1}{2}(\hat{\mathbf{J}}^2 - \hat{\mathbf{L}}^2 - \hat{\mathbf{S}}^2) = \frac{1}{2}(j(j+1) - \ell(\ell+1) - s(s+1))$$

$$\hat{H}_2 = \frac{1}{2m^2c^2} \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r^3} \hat{\mathbf{L}} \cdot \hat{\mathbf{S}}$$

Combining spin 1/2 with angular momentum ℓ, total angular momentum can take values j = ℓ ± 1/2. Corresponding basis states |j = ℓ ± 1/2, m_i, ℓ⟩ diagonalize operator,

$$\hat{\mathbf{L}} \cdot \hat{\mathbf{S}} | j = \ell \pm 1/2, m_j, \ell \rangle = \frac{\hbar^2}{2} \begin{pmatrix} \ell \\ -\ell - 1 \end{pmatrix} | \ell \pm 1/2, m_j, \ell \rangle$$

• Once again, off-diagonal matrix elements of \hat{H}_2 vanish allowing correction to be computed in first order perturbation theory.

$$\langle H_2 \rangle_{n,j=\ell \pm 1/2,m_j,\ell} = \frac{1}{2m^2c^2} \frac{\hbar^2}{2} \begin{pmatrix} \ell \\ -\ell - 1 \end{pmatrix} \frac{Ze^2}{4\pi\epsilon_0} \left\langle \frac{1}{r^3} \right\rangle_{n\ell}$$

$$\langle H_2 \rangle_{n,j=\ell \pm 1/2,m_j,\ell} = \frac{1}{2m^2c^2} \frac{\hbar^2}{2} \begin{pmatrix} \ell \\ -\ell - 1 \end{pmatrix} \frac{Ze^2}{4\pi\epsilon_0} \left\langle \frac{1}{r^3} \right\rangle_{n\ell}$$

• Making use of identity,

$$\left\langle \frac{1}{r^3} \right\rangle_{n\ell} = \left(\frac{mc\alpha Z}{\hbar n} \right)^3 \frac{1}{\ell(\ell+1/2)(\ell+1)}, \qquad \ell > 0$$

$$\langle \hat{H}_2 \rangle_{n,j=\ell \pm 1/2,m_j,\ell} = \frac{1}{4}mc^2 \left(\frac{Z\alpha}{n}\right)^4 \frac{n}{j+1/2} \begin{cases} \frac{1}{j} & j=\ell+1/2\\ -\frac{1}{j+1} & j=\ell-1/2 \end{cases}$$

• Rewriting expression for $\langle \hat{H}_1 \rangle$ in new basis $|n,j=\ell\pm 1/2,m_j,\ell
angle$,

$$\langle \hat{H}_1 \rangle_{n,j=\ell \pm 1/2,m_j,\ell} = -\frac{1}{2}mc^2 \left(\frac{Z\alpha}{n}\right)^4 n \begin{cases} \frac{1}{j} & j=\ell+1/2\\ \frac{1}{j+1} & j=\ell-1/2 \end{cases}$$

• Combining these expressions, for $\ell > 0$, we have

(c) Darwin term

Final relativistic correction arises from "Zitterbewegung" of electron

 giggling – which smears effective potential felt by electron,

$$\hat{H}_{3} = \frac{\hbar^{2}}{8m^{2}c^{2}}\nabla^{2}V = \frac{\hbar^{2}}{8m^{2}c^{2}}\frac{eQ_{\text{nuclear}}(\mathbf{r})}{\epsilon_{0}} = \frac{\pi\hbar^{2}}{2m^{2}c^{2}}\frac{Ze^{2}}{4\pi\epsilon_{0}}\delta^{(3)}(\mathbf{r})$$

• Since perturbation acts only at origin, it effects only $\ell = 0$ states,

$$\langle \hat{H}_3 \rangle_{nj=1/2, m_j \ell=0} = \frac{\pi \hbar^2}{2m^2 c^2} \frac{Ze^2}{4\pi\epsilon_0} |\psi_{n00}(0)|^2 = \frac{1}{2}mc^2 \frac{(Z\alpha)^4}{n^3}$$

• This term is formally identical to that which would be obtained from $\langle \hat{H}_2 \rangle$ at $\ell = 0$. As a result, combining all three contributions,

$$\Delta E_{n,j=\ell\pm 1/2,m_j,\ell} = \frac{1}{2}mc^2\left(\frac{\alpha Z}{n}\right)^4\left(\frac{3}{4}-\frac{n}{j+1/2}\right)$$

independent of ℓ and m_j .

Spectroscopic notation

- To discuss energy shifts for particular states, it is helpful to introduce some nomenclature from atomic physics.
- For a state with principal quantum number n, total spin s, orbital angular momentum ℓ, and total angular momentum j, one may define the state by the spectroscopic notation,

• For a hydrogen-like atom, with just a single electron, 2s + 1 = 2. In this case, the factor 2s + 1 is often just dropped for brevity.

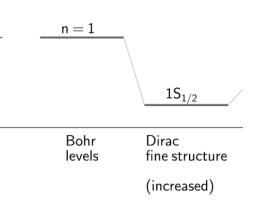
Relativistic corrections

$$\Delta E_{n,j=\ell\pm 1/2,m_j,\ell} = \frac{1}{2}mc^2\left(\frac{\alpha Z}{n}\right)^4\left(\frac{3}{4}-\frac{n}{j+1/2}\right)$$

- For a given n, relativistic corrections depend only on j and n.
- For n = 1, $\ell = 0$ and j = 1/2: Both $1S_{1/2}$ states, with $m_j = \pm 1/2$, experience negative energy shift of $-\frac{1}{4}Z^4\alpha^2$ Ry.
- For n = 2, $\ell = 0, 1$: With j = 1/2, both $2S_{1/2}$ and $2P_{1/2}$ states have shift, $-\frac{5}{64}Z^4\alpha^2 \operatorname{Ry}$, while $2P_{3/2}$ experiences a shift $-\frac{1}{64}Z^4\alpha^2 \operatorname{Ry}$.

 $n = \infty \text{ (Continuum)}$ n = 3 $3P_{3/2}, 3D_{3/2}$ $3D_{5/2}$ $3S_{1/2}, 3P_{1/2}$ n = 2 $2P_{3/2}$ $2S_{1/2}, 2P_{1/2}$

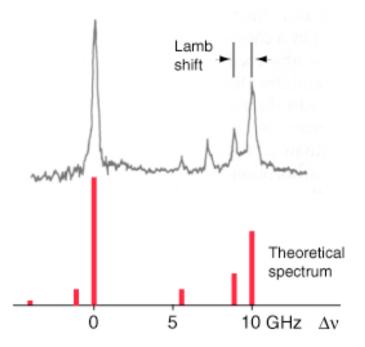




(Further) relativistic corrections: Lamb shift

- Perturbative corrections predicted by Dirac theory predict that, for hydrogen, the $2S_{1/2}$ and $2P_{1/2}$ states should remain degenerate.
- However, in 1951, an experimental study by Willis Lamb discovered that $2P_{1/2}$ state is slightly lower than the $2S_{1/2}$ state Lamb shift.

 Might seem that such a tiny effect would be insignificant, but shift provided considerable insight into quantum electrodynamics.



Lamb shift

- Within framework of quantum electrodynamics, Coulomb interaction is mediated by exchange of photons – "gauge particles".
- Interaction of electron with electromagnetic field can induce a "self-interaction" ~> effective smearing of electron position,

$$\langle (\delta \mathbf{r})^2 \rangle \simeq \frac{2\alpha}{\pi} \left(\frac{\hbar}{mc} \right)^2 \ln \frac{1}{\alpha Z}, \qquad \delta r \sim 10^{-5} a_0$$

- Causes electron spin *g*-factor to be slightly different from 2.
- There is also a slight weakening of the force on the electron when it is very close to the nucleus, causing $2S_{1/2}$ state to be slightly higher in energy than the $2P_{1/2}$ state.

$$\Delta E_{\text{Lamb}} \simeq \frac{1}{2} mc^2 \left(\frac{\alpha Z}{n}\right)^4 n \times \left(\frac{8}{3\pi} \alpha \ln \frac{1}{\alpha Z}\right) \delta_{\ell,0}$$

Hyperfine structure

 Finally, we should address the potential influence of the nuclear spin, I, which leads to a nuclear magnetic moment,

$$\mathbf{M} = g_N \frac{e}{2M_N} \mathbf{I}$$

where nucleus has mass M_N and gyromagnetic ratio g_N .

- Since nucleus has internal structure, g_N is not simply 2. For proton, sole nuclear constituent of atomic hydrogen, $g_p \approx 5.56$. Even though neutron is charge neutral, $g_n \approx -3.83$.
- Magnetic moment generates vector potential $\mathbf{A} = -\frac{\mu_0}{4\pi} \mathbf{M} \times \nabla(1/r)$ and magnetic field

$$\mathbf{B} = \nabla \times \mathbf{A} = \frac{\mu_0}{4\pi} \left[\frac{3\mathbf{r}(\mathbf{r} \cdot \mathbf{M}) - r^2 \mathbf{M}}{r^5} + \frac{8\pi}{3} \mathbf{M} \delta^{(3)}(\mathbf{r}) \right]$$

Hyperfine interaction

$$\mathbf{B} = \nabla \times \mathbf{A} = \frac{\mu_0}{4\pi} \left[\frac{3\mathbf{r}(\mathbf{r} \cdot \mathbf{M}) - r^2 \mathbf{M}}{r^5} + \frac{8\pi}{3} \mathbf{M} \delta^{(3)}(\mathbf{r}) \right]$$

 As a result, we obtain hyperfine interaction with orbital and spin degrees of freedom of electron,

$$\hat{H}_{hyp} = rac{e}{2m}(\hat{\mathbf{L}} + 2\hat{\mathbf{S}}) \cdot \mathbf{B}$$

• Energy level shift of the ground state can be estimated using perturbation theory. If we consider (for simplicity) just the $\ell = 0$ states, only last term in **B** contributes at lowest order, and leads to

$$\langle \hat{H}_{\rm hyp} \rangle_{n,1/2,0} = \frac{\mu_0}{4\pi} \frac{g_N e}{2M_N} \frac{e}{m} \frac{8\pi}{3} |\psi_{n00}(0)|^2 \hat{\mathbf{S}} \cdot \hat{\mathbf{I}} / \hbar^2$$

Hyperfine interaction

$$\langle \hat{H}_{\rm hyp} \rangle_{n,1/2,0} = \frac{\mu_0}{4\pi} \frac{g_N e}{2M_N} \frac{e}{m} \frac{8\pi}{3} |\psi_{n00}(0)|^2 \hat{\mathbf{S}} \cdot \hat{\mathbf{I}} / \hbar^2$$

• With
$$|\psi_{n00}(0)|^2 = rac{1}{\pi n^3} (rac{Zlpha mc}{\hbar})^3$$
, we obtain

$$\langle \hat{H}_{\rm hyp} \rangle_{n,1/2,0} = \frac{1}{2}mc^2 \left(\frac{Z\alpha}{n}\right)^4 n \times \frac{8}{3}g_N \frac{m}{M_N} \hat{\mathbf{S}} \cdot \hat{\mathbf{I}}/\hbar^2$$

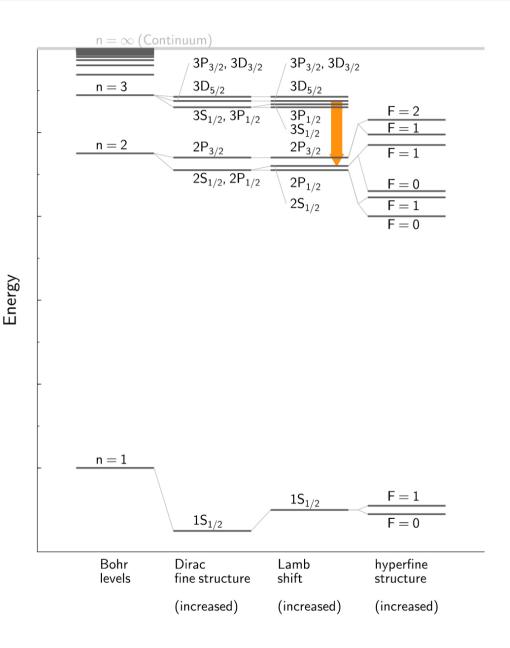
showing scale of perturbation suppressed over fine structure by factor $m/M_N \sim 10^{-3}$.

• Finally, as with spin-orbit interaction, if we set $\mathbf{F} = \mathbf{I} + \mathbf{S}$,

$$\frac{1}{\hbar^2} \mathbf{S} \cdot \mathbf{I} = \frac{1}{2\hbar^2} (\mathbf{F}^2 - \mathbf{S}^2 - \mathbf{I}^2) = \frac{1}{2} (F(F+1) - 3/4 - I(I+1))$$
$$= \frac{1}{2} \begin{cases} I & F = I + 1/2 \\ -I - 1 & F = I - 1/2 \end{cases}$$

Summary of atomic energy scales

- Gross structure: Dictated by orbital kinetic and potential energies, ca. 1 10 eV.
- Fine structure: Relativistic corrections (spin-orbit, etc.) split degenerate multiplets leading to small shift in energy, ca. 10⁻⁴ 10⁻⁵ eV.
- Hyperfine structure: Interaction of electron magnetic moment with field generated by nuclear spin leads to further splitting of multiplets, ca. 10⁻⁷ – 10⁻⁸ eV



	GROU	,	P	ERI	OD		TA	BL	E (DF	TH	ΕE	ELE	M	EN	TS		
	1 6								797	F42/3			http	://www.ktf-	split.hr/peri	odni/en/		18 VIIIA
g	1 1.003	9		RELATI	VE ATOMIC N	AASS (1)	Me		Semimetal	Nonm								2 4.0026
PERIOD	I H		GRO	OUP IUPAC	G	ROUP CAS					ogens elemen	1						He
A.	HYDROGE		1	13 IIIA			Alk	Alkaline earth metal		17 Halogens element		1	13 NIIA	14 IVA	15 VA	16 VIA	17 VIIA	
	3 6.9		ATOMIC 1	NUMBER	5 10.811		Tre	insition metal	5	18 Noble	gas		5 10.811	6 12.011	7 14.007	8 15.999	9 18.998	10 20,180
	² Li	Be		SYMBOL -	B			Lanthanide			(25 °C; 101	kPa)	B	С	N	0	F	Ne
				Z.	BORON			Actinide		- gas - liquid	Fe - solid	tic	BORON	CARBON	NITROGEN	OXYGEN	FLUORINE	NEON
	11 22.9			ELE	MENT NAME					and size		-	13 26,982	14 28.086	15 30.974	16 32.065	17 35.453	18 39.948
	³ Na	Mg				Æ					~ 4		Al	Si	Р	S	CI	Ar
			A CONTRACTOR OF A CONTRACTOR A CO	4 IVB	5 VB	6 VIB	A(Ct	Ire	10	4_		ALUMINIUM	SILICON	PHOSPHORUS	SULPHUR	CHLORINE	ARGON
	19 39.09	0	~	22 47.867	23 50.942	24 51.990		1000	0	20 51.093	63.54		31 69.723	32 72.64	33 74.922	34 78.96	35 79.904	36 83.80
	4 K	Ca	Sc	11	V	Cr	Ivin	Fe	Co	INI	Cu	Zn	Ga	Ge	As	Se	Br	Kr
	37 85.46		SCANDIUM 39 88.906	40 91.224	VANADIUM	CHROMIUM 42 95.94	MANGANESE	IRON 44 101.07	COBALT	NICKEL	COPPER	ZINC 48 442 44	GALLIUM	GERMANIUM 50 118.71	ARSENIC 51 121.76	SELENIUM 52 127.60	8ROMINE	54 131.29
			39 88.906														33 120.90	
	IND	Sr	Y	Zr	- R/	IVLO	- ^{AC}	ele	ct	ror	Ag	to	ma	Sn	Sb	Te	1	Xe
	RUBIDIU			ZIRCONIUM 72 178,49	NICER	74 183 84	75 186 21		CL		79 106 07	80.200 59	11 DILM	TIN 82 207.2	ANTIMONY 83 208.98	TELLURIUM	IODINE 85 (210)	XENON 86 (222)
	c a		57-71 La-Lu		To	W	Do	10 150,25	Tree	D4	1.5 150.57	UT	1204,50					
	0	Ba	Lanthanide	Hf	Ta	new manual	Re	Os	lr	Pt	Au	Hg	11	Pb	Bi	Po	At	Rn
	CAESIUN 87 (22		89-103	HAFNIUM	TANTALUM	TUNGSTEN	RHENIUM 107 (264)	OSMIUM 108 (277)	IRIDIUM 109 (268)	PLATINUM 110 (281)	GOLD	MERCURY 112 (285)	THALLIUM	LEAD	BISMUTH			RADON
	7 Fr	Ra	Ac-Lr	Rf	IDb		IBIh	IEIs	Mit	Uum		ITImlk						
	FRANCIU		Actinide	INTHERFORMUN		Sg								DIIIU				
	A		1			achiertatum								Contractor Data				
				LANTHANIDE Copyright © 1998-2003														
11		e Appl. Chem., 73, No. 4, 667-683 (2001) ative atomic mass is shown with five ifficant figures. For elements have no stable lides, the value enclosed in brackets		57 138.91	58 140.12	59 140.91	60 144.24	61 (145)	62 150.36	63 151.96	64 157.25	65 158.93	66 162.50	67 164.93	68 167.26	69 168.93	70 173.04	71 174.97
	uclides, the v			La	Ce	Pr	Nd	1Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu
indicates the mass number of the longest-lived isotope of the element. However three such elements (Th. Pa, and U)		LANTHANUM	CERIUM		NEODYMIUM	PROMETHIUM	SAMARIUM	EUROPIUM	GADOLINIUM	TERBIUM		HOLMIUM	ERBIUM	THULIUM	YTTERBIUM	LUTETIUM		
However intree such elements (1n, Pa, and 0) do have a characteristic terrestrial isolopic composition, and for these an atomic weight is 89 (227) 90 232.04					91 221 04	97 229 02	03 /2271	94 (244)	95 (242)	96 (247)	97 (247)	98 (251)	99 (252)	100 (257)	101 (258)	102 (259)	103 (262)	
tabulated.			and a second			72 230.03		-				The Const			-			
			Ac	Th	Pa	U	$\mathbb{N}\mathbb{P}$	Pu	Am	Cm	IBIK	CI	IEs	Fun	Midi	No	Lr	
I	iditor: Aditya Va		(Ilinx.com)	ACTINIUM				NEPTUNIUM							FERMIUM			LAWRENGIUM

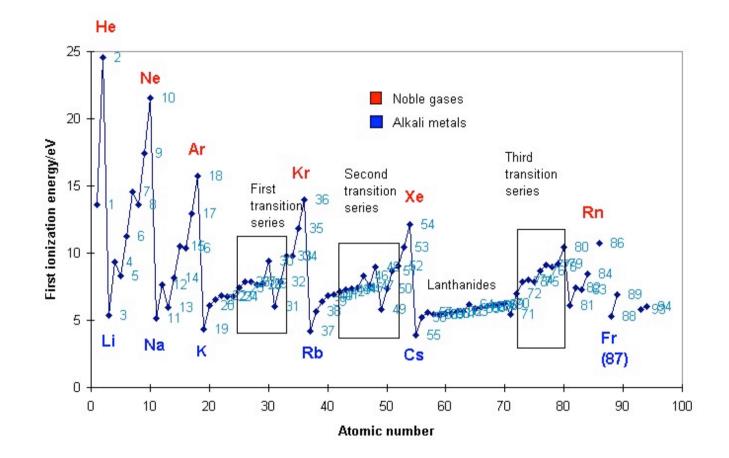
How can we determine energy levels of a multi-electron atom?

- We could start with hydrogenic energy levels for atom of nuclear charge Z, and start filling electrons from lowest levels, accounting for Pauli exclusion.
- Degeneracy for quantum numbers (n, ℓ) is $2 \times (2\ell + 1)$. Each energy level, n, accommodates $2 \times n^2$ electrons:

п	ℓ	Degeneracy in shell	Cumulative total
1	0	2	2
2	0,1	(1+3) imes 2 = 8	10
3	0, 1, 2	(1+3+5) imes 2=18	28
4	0, 1, 2, 3	$(1+3+5+7) \times 2 = 32$	60

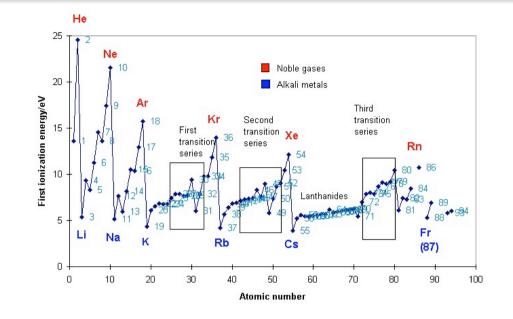
 Expect atoms containing 2, 10, 28 or 60 electrons would be especially stable and that, in atoms containing one more electron, outermost electron would be less tightly bound.

Background: ionization energies of elements



 Instead, find noble gases (Z = 2, 10, 18, 36...) are especially stable, and elements containing one more electron (alkali metals) significantly less tightly bound.

Background



- Failure to predict stable electron configurations reflects omission of electron-electron interaction (cf. our discussion of helium).
- In fact, first ionization energies of atoms show only a weak dependence on Z – outermost electrons are almost completely shielded from nuclear charge:

Effective nuclear charge varies as $Z_{\rm eff} \sim (1 + \gamma)^Z$ where $\gamma > 0$ characterizes "ineffectiveness of screening"; i.e. ionization energy $I_Z = -E_Z \sim Z_{\rm eff}^2 \sim (1 + 2\gamma Z)$ (cf. experiment).

Multi-electron atoms

 Leaving aside (for now) relativistic effects, Hamiltonian for multi-electron atom given by

$$\hat{H} = \sum_{i} \left[-\frac{\hbar^2}{2m} \nabla_i^2 - \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r_i} \right] + \sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}}$$

where $r_{ij} \equiv |\mathbf{r}_i - \mathbf{r}_j|$.

 In addition to nuclear binding potential, there is a further Coulomb interaction between electrons.

As we have seen with helium, this contribution can have important consequences on spectra and spin structure of wavefunction.

 However, electron-electron interaction makes problem "many-body" in character and analytically intractable – we must develop some approximation scheme (even though effects may not be small!).

Multi-electron atoms: outline

- Central field approximation
- Self-consistent field method Hartree approximation
- Structure of the periodic table
- Coupling schemes:
 - LS coupling and Hund's rules
 - 2 jj coupling
- Atomic spectra: selection rules
- Zeeman effect revisited

Central field approximation

• Electron interaction contains large spherically symmetric component arising from "core electrons". Since

$$\sum_{m=-\ell}^{\ell} |Y_{lm}(\theta,\phi)|^2 = \text{const.}$$

closed shell has spherically symmetric charge distribution.

• This suggests a "partitioning" of the Hamiltonian, $\hat{H} = \hat{H}_0 + \hat{H}_1$, with

$$\hat{H}_0 = \sum_i \left[-\frac{\hbar^2}{2m} \nabla_i^2 - \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r_i} + U_i(r_i) \right]$$
$$\hat{H}_1 = \sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}} - \sum_i U_i(r_i)$$

where the radially-symmetric "single-electron potentials", $U_i(r)$, accommodate "average effect" of other electrons, i.e. \hat{H}_1 is small.

Central field approximation

$$\hat{H}_0 = \sum_i \left[-\frac{\hbar^2}{2m} \nabla_i^2 - \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r_i} + U_i(r_i) \right]$$
$$\hat{H}_1 = \sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}} - \sum_i U_i(r_i)$$

- Since single-particle Hamiltonian \hat{H}_0 continues to commute with the angular momentum operator, $[\hat{H}_0, \hat{\mathbf{L}}] = 0$, its eigenfunctions remain indexed by quantum numbers (n, ℓ, m_ℓ, m_s) .
- However, since effective potential, $V(r) + U_i(r)$, is no longer Coulomb-like, ℓ values for a given *n* need not be degenerate.
- But how do we fix U_i(r); the potential energy experienced by each electron depends on the wavefunction of all the other electrons, which is only known after the Schrödinger equation has been solved.
- This suggests an **iterative approach** to solving the problem.

Before embarking on this programme, we should first consider our ambitions:

• The development of computation schemes to address quantum mechanics of many-particle systems is a specialist (and challenging) topic common to physics and chemistry.



- Our interest here is merely in the **outcome** of such investigations, and their ramifications for atomic physics.
- We will therefore discuss (general) principles of the methodology, but the detailed technical aspects of the approach need not be committed to memory!

 To understand how the potentials U_i(r) can be estimated, we will follow a variational approach due to Hartree:

If electrons are (for now) considered distinguishable, wavefunction can be factorized into (normalized) product state,

 $\Psi(\{\mathbf{r}_i\}) = \psi_{i_1}(\mathbf{r}_1)\psi_{i_2}(\mathbf{r}_2)\cdots\psi_{i_N}(\mathbf{r}_N)$

where the quantum numbers, $i \equiv n \ell m_{\ell} m_s$, index individual state occupancies.

- Note that Ψ({r_i}) is not a properly antisymmetrized Slater determinant – exclusion principle taken into account only insofar that we have assigned different quantum numbers, nℓm_ℓm_s.
- In this approximation, if $U_i(r) = 0$, the ground state would involve filling the lowest shells with electrons.

$$\Psi(\{\mathbf{r}_i\}) = \psi_{i_1}(\mathbf{r}_1)\psi_{i_2}(\mathbf{r}_2)\cdots\psi_{i_N}(\mathbf{r}_N)$$

• Variational ground state energy:

$$\Xi = \langle \Psi | \hat{H} | \Psi \rangle = \sum_{i} \int d^{3}r \, \psi_{i}^{*} \left(-\frac{\hbar^{2} \nabla^{2}}{2m} - \frac{1}{4\pi\epsilon_{0}} \frac{Ze^{2}}{r} \right) \psi_{i}$$
$$+ \frac{1}{4\pi\epsilon_{0}} \sum_{i < j} \int d^{3}r \, \int d^{3}r' \, \psi_{i}^{*}(\mathbf{r}) \psi_{j}^{*}(\mathbf{r}') \frac{e^{2}}{|\mathbf{r} - \mathbf{r}'|} \psi_{j}(\mathbf{r}') \psi_{i}(\mathbf{r})$$

- According to variational principle, we must minimize energy by varying $E[\{\psi_i\}]$, subject to normalization condition, $\langle \psi_i | \psi_i \rangle = 1$.
- Latter imposed by set of Lagrange multipliers, ε_i ,

$$\frac{\delta}{\delta\psi_i^*}\left[E-\varepsilon_i\left(\int d^3r|\psi_i(\mathbf{r})|^2-1\right)\right]=0$$

$$\frac{\delta}{\delta\psi_i^*}\left[E-\varepsilon_i\left(\int d^3r|\psi_i(\mathbf{r})|^2-1\right)\right]=0$$

• Following variation, obtain Hartree equations,

$$\left(-\frac{\hbar^2 \nabla^2}{2m} - \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r}\right) \psi_i + \frac{1}{4\pi\epsilon_0} \sum_{j\neq i} \int d^3r' |\psi_j(\mathbf{r}')|^2 \frac{e^2}{|\mathbf{r} - \mathbf{r}'|} \psi_i(\mathbf{r})$$
$$= \varepsilon_i \psi_i(\mathbf{r})$$

• Amongst all possible trial functions ψ_i , set that minimizes energy determined by effective potential,

$$U_i(\mathbf{r}) = rac{1}{4\pi\epsilon_0}\sum_{j
eq i}\int d^3r'\,|\psi_j(\mathbf{r}')|^2rac{e^2}{|\mathbf{r}-\mathbf{r}'|}$$

• To simplify procedure, useful to engineer radial symmetry by replacing $U_i(\mathbf{r})$ by spherical average, $U_i(r) = \int \frac{d\Omega}{4\pi} U_i(\mathbf{r})$.

$$\left(-\frac{\hbar^2 \nabla^2}{2m} - \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r}\right) \psi_i + \frac{1}{4\pi\epsilon_0} \sum_{j\neq i} \int d^3r' \, |\psi_j(\mathbf{r}')|^2 \frac{e^2}{|\mathbf{r} - \mathbf{r}'|} \psi_i(\mathbf{r})$$
$$= \varepsilon_i \psi_i(\mathbf{r})$$

• To fix Lagrange multipliers, ε_i , we can multiply Hartree equations by $\psi_i^*(\mathbf{r})$ and integrate,

$$\epsilon_{i} = \int d^{3}r \,\psi_{i}^{*} \left(-\frac{\hbar^{2}\nabla^{2}}{2m} - \frac{1}{4\pi\epsilon_{0}} \frac{Ze^{2}}{r} \right) \psi_{i}$$
$$+ \frac{1}{4\pi\epsilon_{0}} \sum_{j \neq i} \int d^{3}r' \,d^{3}r \,|\psi_{j}(\mathbf{r}')|^{2} \frac{e^{2}}{|\mathbf{r} - \mathbf{r}'|} |\psi_{i}(\mathbf{r})|^{2}$$

• From this result, we find

$$E = \sum_{i} \epsilon_{i} - \frac{1}{4\pi\epsilon_{0}} \sum_{i < j} \int d^{3}r' \, d^{3}r \, |\psi_{j}(\mathbf{r}')|^{2} \frac{e^{2}}{|\mathbf{r} - \mathbf{r}'|} |\psi_{i}(\mathbf{r})|^{2}$$

 In summary, within the Hartree framework, the multi-electron Hamiltonian is replaced by the effective single-particle Hamiltonian,

$$\hat{H}_0 = \sum_i \left[-\frac{\hbar^2}{2m} \nabla_i^2 - \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r_i} + U_i(r_i) \right]$$

where the central potentials U_i depend self-consistently on the single-particle wavefunctions,

$$U_i(\mathbf{r}) = \int \frac{d\Omega}{4\pi} \frac{1}{4\pi\epsilon_0} \sum_{j\neq i} \int d^3 \mathbf{r}' \, |\psi_j(\mathbf{r}')|^2 \frac{e^2}{|\mathbf{r} - \mathbf{r}'|}$$

 Once U_is are found, perturbation theory can be applied to residual Coulomb interaction,

$$\hat{H}_1 = \sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}} - \sum_i U_i(r_i)$$

Hartree-Fock method

• An improvement on this procedure can be achieved by introducting a trial variational state wavefunction involving a Slater determinant,

$$\Psi = \frac{1}{\sqrt{N!}} \begin{vmatrix} \psi_1(\mathbf{r}_1) & \psi_1(\mathbf{r}_2) & \psi_1(\mathbf{r}_3) & \cdots \\ \psi_2(\mathbf{r}_1) & \psi_2(\mathbf{r}_2) & \psi_2(\mathbf{r}_3) & \cdots \\ \psi_3(\mathbf{r}_1) & \psi_3(\mathbf{r}_2) & \psi_3(\mathbf{r}_3) & \cdots \\ \vdots & \vdots & \vdots & \ddots \end{vmatrix}$$

where $\psi_k(\mathbf{r}_i)$, with $i = 1, 2 \cdots N$, denote the single-particle wavefunctions for electron *i*, and $k = (n \ell m_\ell m_s)$

• A variational analysis leads to Hartree-Fock equations with additional exchange contribution,

$$\varepsilon_{i}\psi_{i}(\mathbf{r}) = \left[-\frac{\hbar^{2}}{2m}\nabla_{i}^{2} - \frac{1}{4\pi\epsilon_{0}}\frac{Ze^{2}}{r_{i}}\right]\psi_{i}(\mathbf{r})$$
$$+\sum_{j\neq i}\int d^{3}r_{j}^{\prime}\frac{1}{4\pi\epsilon_{0}}\frac{e^{2}}{|\mathbf{r}-\mathbf{r}^{\prime}|}\psi_{j}^{*}(\mathbf{r}^{\prime})\left[\psi_{j}(\mathbf{r}^{\prime})\psi_{i}(\mathbf{r}) - \psi_{j}(\mathbf{r})\psi_{i}(\mathbf{r}^{\prime})\delta_{m_{s_{i}},m_{s_{j}}}\right]$$

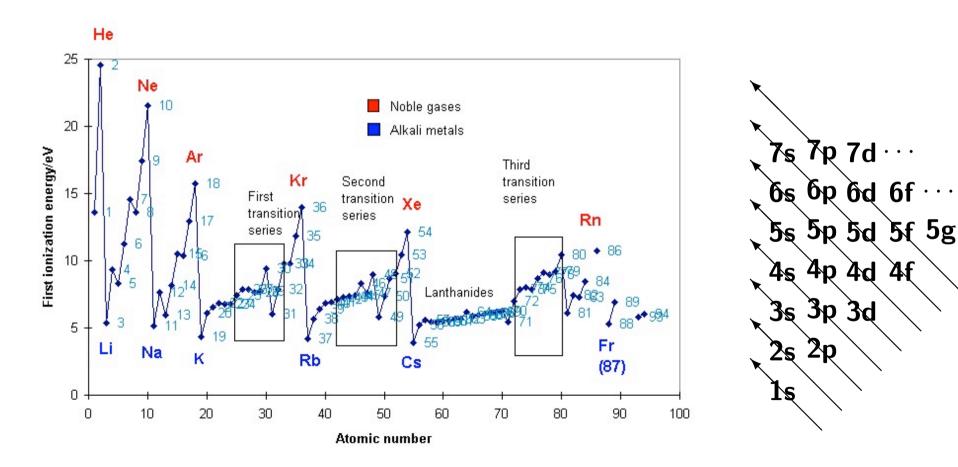
Central field approximation: conclusions

- Although states characterized by quantum numbers nlm_lm_s, Hartree-Fock calculations show that those with different l for given n are now non-degenerate – large l values more effectively screened and lie higher in energy.
- States corresponding to particular *n* referred to as a shell, and those belonging to *n*, ℓ are a subshell. Energy levels ordered as

Subshell name	1 <i>s</i>	2 <i>s</i>	2p	3 <i>s</i>	3 <i>p</i>	4 <i>s</i>	3 <i>d</i>	4 <i>p</i>	5 <i>s</i>	4 <i>d</i>	•••
n =	1	2	2	3	3	4	3	4	5	4	•••
$\ell =$	0	0	1	0	1	0	2	1	0	2	•••
Degeneracy	2	2	6	2	6	2	10	6	2	10	•••
Cumulative	2	4	10	12	18	20	30	36	38	48	•••

Central field approximation: conclusions

Subshell name *s s p d s s* 3p *s p d* . . . Cumulative . . .



- Can use energy sequence to predict ground state electron configuration – fill levels accounting for exclusion aufbau principle.
- Sadly, there are exceptions to rule: e.g. Cu (Z = 29) expected to have configuration (Ar)(4s)²(3d)⁹, actually has (Ar)(4s)¹(3d)¹⁰.

1	Н	(1s)	13.6
2	He	$(1s)^2$	24.6
3	Li	He (2 <i>s</i>)	5.4
4	Be	He (2 <i>s</i>) ²	9.3
5	В	He $(2s)^2(2p)$	8.3
6	C	He $(2s)^2(2p)^2$	11.3
7	Ν	He $(2s)^2(2p)^3$	14.5
8	0	He $(2s)^2(2p)^4$	13.6
9	F	He $(2s)^2(2p)^5$	17.4
10	Ne	He $(2s)^2(2p)^6$	21.6
11	Na	Ne (3 <i>s</i>)	5.1
12	Mg	Ne (3 <i>s</i>) ²	7.6
14	Si	Ne $(3s)^2(3p)^2$	8.1
16	S	Ne $(3s)^2(3p)^4$	10.4
18	Ar	Ne $(3s)^2(3p)^6$	15.8
19	K	Ar (4 <i>s</i>)	4.3

Periodic table

• Aufbau principle forms basis of **Periodic table of elements**: elements with similar electron configurations in outermost shells have similar chemical properties.

	1 1.0079			RELATI	E ATOMIC N	TOMIC MASS (1)		Metal Semimetal Nonmetal										18 VIIA 2 4.0026 He
¹ H			GRC	ROUP IUPAC GROUP CAS			Alk	ali metal		16 Chalcogens element								
h	HYDROGEN	2 11A	ATOMICS	ATOMIC NUMBER 5 10.811				Alkaline earth metal II7 Halogens element				13 IIIA 14 IVA 15 VA 16 VIA 17 VIA 5 10.811 6 12.011 7 14.007 8 15.999 9 18.998						
	3 6.941	4 9.0122					Tra	Transition metals 18 Noble gas			gas		_	6 12.011				10 20.1
	Li	Be	5	SYMBOL -	B		-	Lanthanide			(25 °C; 101)	kPa)	B	C	N	0	F	Ne
-	LITHIUM	BERYLLIUM			BORON			Actinide		- gas - liquid	Fe - solid	tio	BORON	CARBON	NITROGEN	OXYGEN	FLUORINE	NEON
	11 22.990	12 24.305	- /	ELE	MENT NAME				Ja	- indoio	the - synulo		13 26.982	14 28.086	15 30.974	16 32.065	17 35.453	18 39.9
1	Na	Mg							- VIIB -				Al	Si	P	S	Cl	Ar
7	SODIUM	MAGNESIUM	3 IIIB	4 IVB	5 / VB	6 /VIB	7 VIIB	8	9	10	11 18	12 118	ALUMINIUM	SILICON	PHOSPHORUS	SULPHUR	CHLORINE	ARGO
	19 39.098	20 40.078	21 44.956	22 47.867	23 50.942	24 51.996	25 54.938	26 55.845	27 58.933	28 58.693	29 63.546	30 65.39	31 69.723	32 72.64	33 74.922	34 78.96	35 79.904	36 83
1	K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
	POTASSIUM	CALCIUM	SCANDIUM	TITANIUM	VANADIUM	CHROMIUM	MANGANESE	IRON	COBALT	NICKEL	COPPER	ZINC	GALLIUM	GERMANIUM	ARSENIC	SELENIUM	BROMINE	KRYPTO
	37 85.468	38 87.62	39 88.906	40 91.224	41 92.906	42 95.94	43 (98)	44 101.07	45 102.91	46 106.42	47 107.87	48 112.41	49 114.82	50 118.71	51 121.76	52 127.60	53 126.90	54 131
5	Rb	Sr	Y	Zr	Nb	Mo	TC	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Те	I	Xe
	RUBIDIUM	STRONTIUM	YTTRIUM	ZIRCONIUM	NIOBIUM	MOLYBDENUM	TECHNETIUM	RUTHENIUM	RHODIUM	PALLADIUM	SILVER	CADMIUM	INDIUM	TIN	ANTIMONY	TELLURIUM	IODINE	XENO
	55 132.91	56 137.33	57-71	72 178.49	73 180.95	74 183.84	75 186.21	76 190.23	77 192.22	78 195.08	79 196.97	80 200.59	81 204.38	82 207.2	83 208.98	84 (209)	85 (210)	86 (2
5	Cs	Ba	La-Lu	Hf	Та	W	Re	Os	Ir	Pt	Au	Hg	TI	Pb	Bi	Po	At	Rn
	CAESIUM	BARIUM	Lanthanide	HAFNIUM	TANTALUM	TUNGSTEN	RHENIUM	OSMIUM	IRIDIUM	PLATINUM	GOLD	MERCURY	THALLIUM	LEAD	BISMUTH	POLONIUM	ASTATINE	RADO
	87 (223)	88 (226)	89-103	104 (261)	105 (262)	106 (266)	107 (264)	108 (277)	109 (268)	110 (281)	111 (272)	112 (285)		114 (289)				
	Fr	Ra	Ac-Lr	IRf	IDh	Sg	IBIh	IBIs	MItt	Www	เป็นหมาย	Umb		Uua				
	FRANCIUM	RADIUM	Actinide	RUTHERFORDIUM		SEABORGIUM		HASSIUM			UNUNUNIUM			UNUNQUADUN				
/			/	LANTHANI	1								,		,	Copyright © 19	98-2003 EniG. ()	eni@ktf-so
		3. No. 4, 667-68	33 (2001)	57 138.91		59 140.91	60 144.24	61 (145)	62 150.36	63 151.96	64 157.25	65 158.93	66 162.50	67 164.93			70 173.04	
elative atomic mass is shown with five gnificant figures. For elements have no stable		no stable	La	Ce	Pr	Nd	IPm	Sm	Eu	Gd	Tb	Dv	Ho	Er	Tm	Yb	Lu	
dica		e enclosed in umber of the lon		LANTHANUM			~	PROMETHIUM			GADOLINIUM		DYSPROSIUM		ERBIUM	THULIUM	YTTERBIUM	LUTETIL
owe	ver three such	elements (Th, F		ACTINIDE	OLIVON	TOULOUTINU		I TOME THOM	or of the second	Lontor Iom	Condo Ention	TERDION	Did Room	TOLMON	Litolom	morion	TTERBION	LOILING
mp	osition, and for	eristic terrestria these an atomic		89 (227)	90 232.04	91 231.04	92 238.03	93 (237)	94 (244)	95 (243)	96 (247)	97 (247)	98 (251)	99 (252)	100 (257)	101 (258)	102 (259)	103 (2
tabulated.			Ac	Th	Pa	U	NID	Pu	Am	Cm	Bk	Cf	Es	Fm	MId	No	ILn	
		110		1	U	TANK	21 00	X-171001	Cunn	Terns		100 M	T. ma	TATION .	T AM	1001		

Coupling schemes

- The aufbau principle predicts ground state occupation of subshells but does not specify spin and orbital angular momenta of subshells.
- To deal with this question, we must consider spin-orbit and residual Coulomb interaction between outer electrons.
- Hamiltonian for multi-electron atom can be written as,

$$\hat{H} \approx \hat{H}_0 + \underbrace{\sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}} - \sum_i U_i(r)}_{\hat{H}_1} + \underbrace{\sum_i \xi_i(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i}_{\hat{H}_2}$$

where \hat{H}_0 includes central field terms, \hat{H}_1 is residual Coulomb interaction, and \hat{H}_2 is spin-orbit interaction.

Coupling schemes

$$\hat{H} \approx \hat{H}_0 + \underbrace{\sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}} - \sum_i U_i(r)}_{\hat{H}_1} + \underbrace{\sum_i \xi_i(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i}_{\hat{H}_2}$$

- For light atoms, $\hat{H}_1 \gg \hat{H}_2$, can treat \hat{H}_2 as a perturbation on $\hat{H}_0 + \hat{H}_1$ known as LS (or Russell-Saunders) coupling.
- For heavy atoms (or ionized light atoms), $\hat{H}_2 \gg \hat{H}_1$, electrons become relativistic and spin-orbit interaction dominates jj coupling.

Both scenarios are approximations – real atoms do not always conform to this "comparatively simple" picture.

Coupling schemes: LS coupling

$$\hat{H} \approx \hat{H}_{0} + \underbrace{\sum_{i < j} \frac{1}{4\pi\epsilon_{0}} \frac{e^{2}}{r_{ij}} - \sum_{i} U_{i}(r)}_{\hat{H}_{1}} + \underbrace{\sum_{i} \xi_{i}(r_{i})\hat{\mathbf{L}}_{i} \cdot \hat{\mathbf{S}}_{i}}_{\hat{H}_{2}}}_{\hat{H}_{2}}$$

- Since \hat{H} commutes with set of total angular momenta, \hat{J}^2 , \hat{L}^2 , and \hat{S}^2 , energy levels of multi-electron atoms are characterized by quantum numbers L, S, J.
- Their ordering in energy set by **Hund's rules**.
- As rules empirical, there are exceptions. Moreover, as atomic mass increases and electrons become relativistic, spin-orbit interactions become increasingly important further undermining rules.

Coupling schemes: LS coupling and Hund's rules

$$\hat{H} \approx \hat{H}_0 + \underbrace{\sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}} - \sum_i U_i(r)}_{\hat{H}_1} + \underbrace{\sum_i \xi_i(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i}_{\hat{H}_2}$$

Combine spins to obtain possible values of total spin S. (Remember that closed shells contribute zero spin.)

The largest permitted value of *S* lies lowest in energy.

Physically: maximising S makes spin wavefunction as symmetric as possible: tends to make spatial wavefunction antisymmetric, reduces Coulomb repulsion (cf. helium).

Coupling schemes: LS coupling and Hund's rules

$$\hat{H} \approx \hat{H}_0 + \underbrace{\sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}} - \sum_i U_i(r)}_{\hat{H}_1} + \underbrace{\sum_i \xi_i(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i}_{\hat{H}_2}$$

Por each value of S, find the possible values of total angular momentum L. (Remember that closed shells contribute zero orbital angular momentum.)

The largest permitted value of *L* lies lowest in energy.

Physically: maximising L also tends to keep the electrons apart.

 In deciding on permitted values of L and S, we also have to ensure that both quantum statistics and the exclusion principle is respected,
 i.e. total electron wavefunction must be antisymmetric under particle exchange.

Coupling schemes: LS coupling and Hund's rules

$$\hat{H} \approx \hat{H}_{0} + \underbrace{\sum_{i < j} \frac{1}{4\pi\epsilon_{0}} \frac{e^{2}}{r_{ij}} - \sum_{i} U_{i}(r)}_{\hat{H}_{1}} + \underbrace{\sum_{i} \xi_{i}(r_{i})\hat{\mathbf{L}}_{i} \cdot \hat{\mathbf{S}}_{i}}_{\hat{H}_{2}}}_{\hat{H}_{2}}$$

Ouple L and S to obtain values of J (hence name of scheme). (Remember that closed shells contribute zero angular momentum.)

If subshell is less than half full, smallest value of J lies lowest in energy; otherwise, largest value lies lowest.

Energy separation for different J arises from treating spin-orbit term as a perturbation (fine structure),

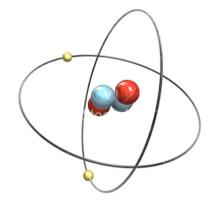
$$\langle Jm_J LS | \sum_i \xi_i(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i | Jm_J LS \rangle = \zeta(L, S) \langle Jm_J LS | \hat{\mathbf{L}} \cdot \hat{\mathbf{S}} | Jm_J LS \rangle$$

= $\zeta(L, S) [J(J+1) - L(L+1) - S(S+1)]/2$

Since sign of $\zeta(L, S)$ changes according to the whether the subshell is more or less than half-filled, the third Hund's rule is established.

LS coupling – Example: helium

Helium has ground state electron configuration (1s)², i.e. L = S = J = 0.
 N.B. For any completely filled subshell, L = S = 0 and hence J = 0.



- For excited state, e.g. $(1s)^1(2p)^1$, can have S = 1 or S = 0, with S = 1 state lying lower in energy according to Hund's rules.
- Combining orbital angular momenta gives L = 1 and, with S = 0, J = 1, while with S = 1, J = 0, 1, 2 with J = 0 lowest in energy.
- In spectroscopic notation ${}^{2S+1}L_J$, four possible states, ${}^{3}P_0 {}^{3}P_1 {}^{3}P_2$ and ${}^{1}P_1$ where three ${}^{3}P$ states separated by spin-orbit interaction, and singlet ${}^{1}P$ state lies much higher in energy due to Coulomb.

Landé interval rule

• Since separation of energies for states of different J arises from spin-orbit term contribution \hat{H}_2 (fine structure),

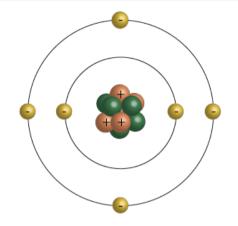
$$egin{aligned} |J,m_J,L,S| &\sum_i \xi_i(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i |J,m_J,L,S
angle \ &= rac{\zeta(L,S)}{2} [J(J+1) - L(L+1) - S(S+1)] \end{aligned}$$

separation between pair of adjacent levels in a fine structure multiplet is proportional to larger of two J values,

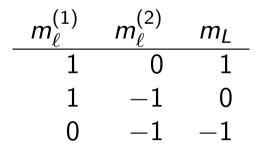
$$\Delta_J \propto J(J+1) - (J-1)J = 2J$$

e.g. separation between ${}^{3}P_{2}$ and ${}^{3}P_{1}$, and ${}^{3}P_{1}$ and ${}^{3}P_{0}$ should be in ratio 2:1.

- Carbon has ground state electron configuration $(1s)^2(2s)^2(2p)^2$.
- With two *identical* electrons in same *unfilled* subshell, wavefunction must be antisymmetric.



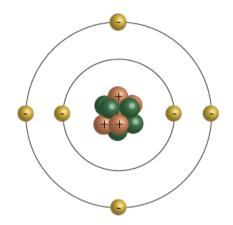
- Total spin can either be singlet S = 0 (antisymmetric) or one of the triplet S = 1 states (symmetric).
- To form antisymmetric total angular momentum state, two electrons must have different values of m_{ℓ}
- Inspecting the values of m_L we can deduce that L = 1.



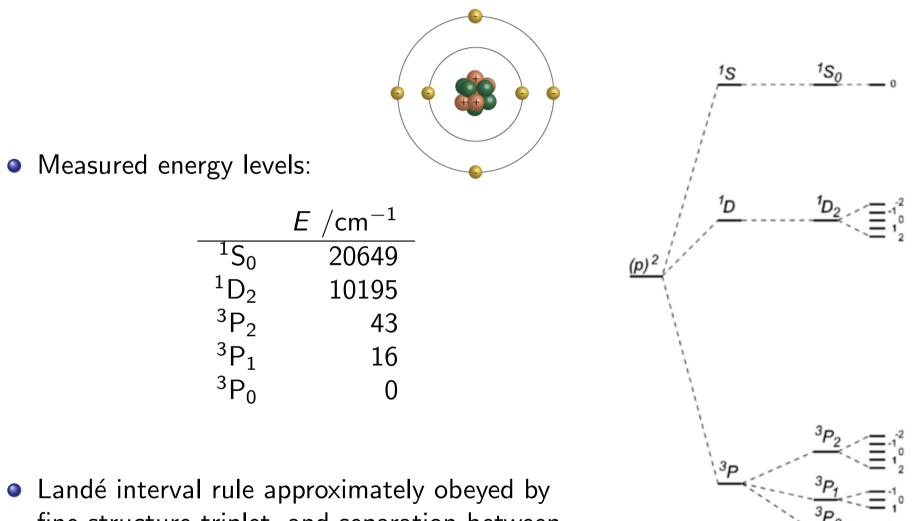
 m_L



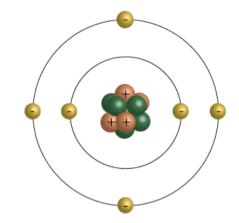
- Carbon has ground state electron configuration $(1s)^2(2s)^2(2p)^2$.
- With two *identical* electrons in same *unfilled* subshell, wavefunction must be antisymmetric.



- To ensure antisymmetry of wavefunction, we must therefore take S = 1 with L = 1 and S = 0 with L = 2 or 0.
- To account for fine structure, states with S = 1 and L = 1 can be combined into single J = 0 state, three J = 1 states, and five J = 2 states leading to terms ${}^{3}P_{0}$, ${}^{3}P_{1}$, and ${}^{3}P_{2}$ respectively.
- Similarly the S = 0, L = 2 state can be combined to give five J = 2 states, ${}^{1}D_{2}$, while S = 0, L = 0 state gives single J = 0 state, ${}^{1}S_{0}$.



fine structure triplet, and separation between L and S values caused by Coulomb repulsion is *much* greater than spin-orbit effect.



- For excited states of carbon, e.g. (2p)¹(3p)¹, electrons are no longer equivalent because they have different radial wavefunctions.
- We can now combine any of S = 0, 1 with any of L = 0, 1, 2, yielding the following terms (in order of increasing energy, according to Hund's rules):

$$^{3}D_{1,2,3}$$
 $^{3}P_{0,1,2}$ $^{3}S_{1}$ $^{1}D_{2}$ $^{1}P_{1}$ $^{1}S_{0}$

- Our studies of the energy spectrum of atomic hydrogen using the non-relativistic Schrödinger equation showed that states are organised in a shell structure, indexed by a principle quantum number n and characterised by an n²-fold degeneracy.
- To address the electronic structure of multielectron atoms, we have to accommodate two classes of additional effects:
 - Even hydrogenic (i.e. single-electron) atoms are subject to corrections from relativistic effects (spin-orbit coupling, etc.) fine structure, vacuum fluctuations of EM field Lamb shift, and interaction with nuclear spin hyperfine structure which together conspire to lift state degeneracy.
 - In addition, in multielectron atoms, the direct Coulomb interaction between electrons lead to screening of the nuclear charge, and rearrange the ordering of the shell structure.

 Although electron-electron interactions make the multielectron system formally intractable, the spherical symmetry of filled core electron states justifies central field approximation in which the principle effect of interactions is captured by a single-particle potential,

$$\hat{H}_{0} = \sum_{i} \left[-\frac{\hbar^{2} \nabla_{i}^{2}}{2m} - \frac{Z e^{2}}{4\pi\epsilon_{0} r_{i}} + U_{i}(r_{i}) \right], \quad \hat{H}_{1} = \sum_{i < j} \frac{e^{2}}{4\pi\epsilon_{0} r_{ij}} - \sum_{i} U_{i}(r_{i})$$

- Numerical studies (based on self-consistent Hartree-Fock scheme) provide a simple phenomenology to describe energy ordering of core subshells – aufbau principle
- Influence of residual electron interaction, \hat{H}_1 , and relativistic spin-orbit corrections

$$\hat{H}_2 = \sum_i \xi(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i$$

on valence states can then be addressed within perturbation theory.

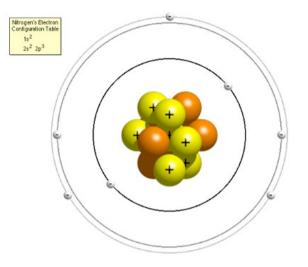
$$\hat{H} \approx \hat{H}_0 + \underbrace{\sum_{i < j} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r_{ij}} - \sum_i U_i(r)}_{\hat{H}_1} + \underbrace{\sum_i \xi_i(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i}_{\hat{H}_2}$$

- For light atoms, $\hat{H}_1 \gg \hat{H}_2$, can treat \hat{H}_2 as a perturbation on $\hat{H}_0 + \hat{H}_1$ known as LS (or Russell-Saunders) coupling.
- For heavy atoms (or ionized light atoms), $\hat{H}_2 \gg \hat{H}_1$, electrons become relativistic and spin-orbit interaction dominates jj coupling.

- In LS coupling, the ground state electron configure is specified by an emperical set of rules known as Hund's rules. Subject to Pauli exclusion:
 - The largest permitted value of total S lies lowest in energy.
 - 2 The largest permitted value of total L lies lowest in energy.
 - If subshell is less than half full, smallest value of total J lies lowest in energy; otherwise largest value lies lowest.

LS coupling – Example: nitrogen

- Nitrogen has ground state electron configuration $(1s)^2(2s)^2(2p)^3$.
- The maximal value of spin is S = 3/2while L can take values 3, 2, 1 and 0.



- Since spin wavefunction (being maximal) is symmetric, spatial wavefunction must be antisymmetric − all three states with m_ℓ = 1, 0, −1 must be involved.
- We must therefore have L = 0 and J = 3/2 with the term, ${}^{4}S_{3/2}$.

jj coupling scheme

$$\hat{H} \approx \hat{H}_{0} + \underbrace{\sum_{i < j} \frac{1}{4\pi\epsilon_{0}} \frac{e^{2}}{r_{ij}} - \sum_{i} U_{i}(r)}_{\hat{H}_{1}} + \underbrace{\sum_{i < j} \hat{\xi}_{i}(r_{i})\hat{\mathbf{L}}_{i} \cdot \hat{\mathbf{S}}_{i}}_{\hat{H}_{2}}}_{\hat{H}_{2}}$$

- When relativistic effects dominate residual electrostatic interaction, *Ĥ*₁, (i.e. heavy elements) electrons move independently in central field, subject to spin-orbit interaction. In this limit, states are both eigenstates of *Ĵ*² (as before), and also of *Ĵ*²_i for each electron.
- In *jj* coupling, separate energy shifts independent of total J and M_J , $\Delta E = \langle n_i \ell_i s_i j_i J m_J | \sum_i \xi(r_i) \hat{\mathbf{L}}_i \cdot \hat{\mathbf{S}}_i | n_i \ell_i s_i j_i J m_J \rangle = \sum_i \Delta E_i$ where

$$\Delta E_i = \zeta(n_i, \ell_i) \left[j_i(j_i + 1) - \ell_i(\ell_i + 1) - s_i(s_i + 1) \right] / 2$$

• The degeneracy with respect to J is then lifted by the small electrostatic interaction between electrons, \hat{H}_1 .

jj coupling scheme: Example

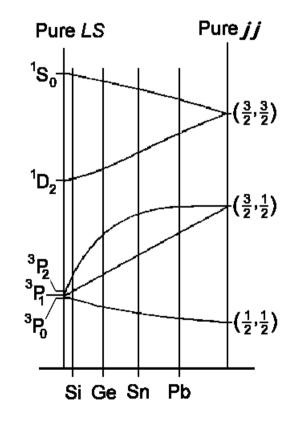
- Consider configuration $(np)^2$ (cf. carbon in LS scheme): Combining s = 1/2 with $\ell = 1$, each electron can have j = 1/2 or 3/2.
- If electrons have same j value, they are equivalent, so we have to take care of symmetry:
 - (a) $j_1 = j_2 = 3/2 \Rightarrow J = 3, 2, 1, 0$, of which J = 2, 0 are antisymmetric.
 - (b) $j_1 = j_2 = 1/2 \Rightarrow J = 1, 0$, of which J = 0 is antisymmetric. (c) $j_1 = 1/2, \ j_2 = 3/2 \Rightarrow J = 2, 1$.
- Taking into account Pauli exclusion, in *jj* coupling (where the term is written $(j_1, j_2)_J$), we have the following terms:

$(1/2, 1/2)_0$ $(3/2, 1/2)_1$ $(3/2, 1/2)_2$ $(3/2, 3/2)_2$ $(3/2, 3/2)_0$

in order of increasing energy.

$(1/2, 1/2)_0$ $(3/2, 1/2)_1$ $(3/2, 1/2)_2$ $(3/2, 3/2)_2$ $(3/2, 3/2)_0$

- Both LS and *jj* coupling give same J values (two states with J = 0, two with J = 2 and one with J = 1) and in same order.
- However, pattern of levels different: in LS coupling we found a triplet (³P₀, ³P₁, ³P₂) and two singlets (¹D₂ and ¹S₀), while in ideal *jj* scenario, we have two doublets and a singlet.
- The sets of states in two schemes must be expressible as linear combinations of one another, and physical states for real atom likely to differ from either approximation – e.g. *jj* coupling not seen in Pb(6*p*)² but is seen in Cr¹⁸⁺ which has same configuration as carbon, (2*p*)².

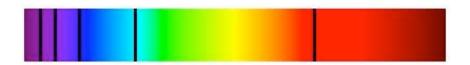


Atomic spectra

- Atomic spectra result from transitions between different electronic states of an atom via emission or absorption of photons.
- In emission spectra, atom is excited by some means (e.g. thermally through collisions), and one observes discrete spectral lines in light emitted as atoms relax.



 In absorption spectra, one illuminates atoms using a broad waveband source, and observes dark absorption lines in the spectrum of transmitted light.



 Atoms excited in this process subsequently decay by emitting photons in random directions – fluorescence.

Atomic spectra: selection rules

Basic theory governing emission and absorption will be outlined in detail when we study radiative transitions. Here we anticipate some results:

• In electric dipole approximation, rate of transitions is proportional to matrix elements of electric dipole operator, $\hat{\mathbf{d}} = -e \sum_{i} \mathbf{r}_{i}$,

$$\Gamma \propto \omega^3 |\langle \psi_{\mathrm{f}} | \hat{\mathbf{d}} | \psi_{\mathrm{i}}
angle |^2, \qquad \omega = |E_{\mathrm{f}} - E_{\mathrm{i}}|$$

- Form of dipole operator, $\hat{\mathbf{d}}$ means that some matrix elements vanish \rightsquigarrow selection rules. For a transition to take place:
 - Parity must change
 - 2 $\Delta J = \pm 1, 0$ (but $0 \rightarrow 0$ is not allowed) and $\Delta M_J = \pm 1, 0$
- Atomic states always eigenstates of parity and $\hat{\mathbf{J}}^2$, so selection rules can be regarded as absolutely valid in electric dipole transitions.

Atomic spectra: selection rules

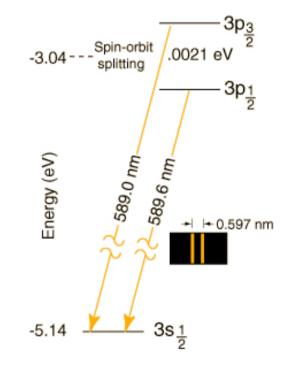
- In specific coupling schemes, further selection rules apply. In the case of ideal LS coupling, we also require:
 - $\Delta S = 0$ and $\Delta M_S = 0$

Follows from conservation of total spin in transition.

- ② $\Delta L = \pm 1$, 0 (but 0 \rightarrow 0 is not allowed) and $\Delta M_L = \pm 1$, 0 Follows from 1. and rules relating to *J*.
- Δℓ_i = ±1 if only electron *i* is involved in transition.
 Follows from parity change rule since the parity of atom is product of parities of separate electron wavefunctions, (-1)^{ℓ_i}.
- However, since LS coupling is only an approximation, these rules should themselves be regarded as approximate.

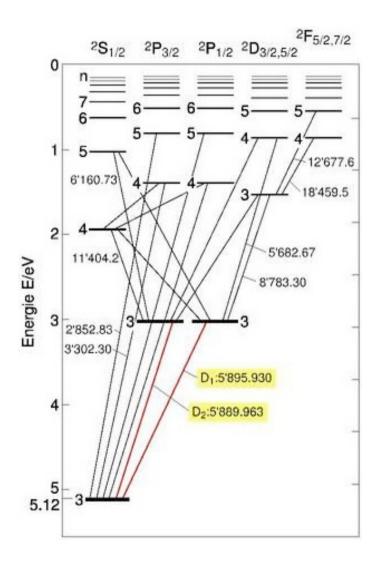
Atomic spectra: single electron atoms

- For "single electron atoms", e.g. alkali metals such as sodium, and also hydrogen, ground state is $(ns)^1$.
- Ground state has term ${}^{2}S_{1/2}$ while excited states all doublets with $J = L \pm 1/2$ (except for *s* states which have J = 1/2).
- Since parity given by (−1)^ℓ, allowed transitions involve Δℓ = ±1, i.e. s ↔ p, p ↔ d, etc. (Larger changes in ℓ contravene ΔJ rule.)
- The s ↔ p transitions are all doublets. In sodium, transition 3s ↔ 3p gives rise to familiar yellow sodium "D-lines" at 589 nm.



Atomic spectra: single electron atoms

- $p \leftrightarrow d$ transitions involve two doublets, ${}^{2}P_{1/2,3/2}$ and ${}^{2}D_{3/2,5/2}$. However, the ${}^{2}P_{1/2} \leftrightarrow {}^{2}D_{5/2}$ transition forbidden by ΔJ rule, so line is actually a triplet.
- As *n* increases, levels approach those for hydrogen, as nuclear charge is increasingly screened by inner electrons.
- In an absorption spectrum, atoms start from ground state, so only ns → n'p lines seen. In emission, atoms are excited into essentially all their excited levels, so many more lines will be seen in the spectrum.



To conclude survey of atomic structure, we now return to consider how atomic spectra are influenced by a magnetic field?

• Begin with hydrogen-like atoms involving just a single electron. In a magnetic field, $\hat{H} = \hat{H}_0 + \hat{H}_{rel.} + \hat{H}_{Zeeman}$, where

$$\hat{H}_{\text{Zeeman}} = -\frac{e}{2mc}B(\hat{L}_z + 2\hat{S}_z) = -\mu_{\text{B}}B(\hat{L}_z + 2\hat{S}_z)/\hbar$$

denotes Zeeman term.

- Since we are dealing with confined electrons, we have neglected the diamagnetic contribution to the Hamiltonian.
- Depending on scale of magnetic field, the spin-orbit term in $\hat{H}_{rel.}$ or the Zeeman term may dominate the spectrum of the atom.

$$\hat{H}_{\text{Zeeman}} = -\frac{e}{2mc}B(\hat{L}_z + 2\hat{S}_z) = -\mu_{\text{B}}B(\hat{L}_z + 2\hat{S}_z)/\hbar$$

 Previously we have seen that, to leading order, relativistic corrections lead to fine-structure energy shift,

$$\Delta E_{n,j}^{\text{rel.}} = \frac{1}{2}mc^2 \left(\frac{Z\alpha}{n}\right)^4 \left(\frac{3}{4} - \frac{n}{j+1/2}\right)$$

for states $|n,j=\ell\pm 1/2,m_j,\ell
angle$.

• For weak magnetic fields, we can also treat Zeeman energy in framework of perturbation theory:

Although states with common j (e.g. $2S_{1/2}$ and $2P_{1/2}$) are degenerate, spatial wavefunctions have different parity, and off-diagonal matrix elements of \hat{H}_{Zeeman} vanish – avoids need for degenerate perturbation theory.

$$\hat{H}_{\text{Zeeman}} = -\frac{e}{2mc}B(\hat{L}_z + 2\hat{S}_z) = -\mu_{\text{B}}B(\hat{L}_z + 2\hat{S}_z)/\hbar = -\mu_{\text{B}}B(\hat{J}_z + \hat{S}_z)/\hbar$$

 Making use of identity, (exercise – refer back to addition of angular momentum and spin)

$$\langle n,j=\ell\pm 1/2,m_j,\ell|\hat{S}_z|n,j=\ell\pm 1/2,m_j,\ell\rangle=\pm\frac{\hbar m_j}{2\ell+1}$$

we obtain the following expression for the first order energy shift,

$$\Delta E_{j=\ell\pm 1,m_j,\ell}^{
m Zeeman} = \ell\pm 1/2, m_j, \ell
angle = \mu_{
m B} B m_j \left(1\pm rac{1}{2\ell+1}
ight)$$

i.e. all degenerate levels split by field.

• In contrast to the "normal" Zeeman effect, the magnitude of the splitting now depends on ℓ .

$$\hat{H}_{\text{Zeeman}} = -\frac{e}{2mc}B(\hat{L}_z + 2\hat{S}_z) = -\mu_{\text{B}}B(\hat{L}_z + 2\hat{S}_z)/\hbar = -\mu_{\text{B}}B(\hat{J}_z + \hat{S}_z)/\hbar$$

- For multi-electron atom in weak field, unperturbed states given by $|J, M_J, L, S\rangle$, where J, L, S refer to total angular momenta.
- To determine Zeeman shift, need to determine $\langle \hat{S}_z \rangle$, presenting an opportunity uto revise angular momenta:
- **1** First we note that the operator $2\hat{\mathbf{L}} \cdot \hat{\mathbf{S}} = \hat{\mathbf{J}}^2 \hat{\mathbf{L}}^2 \hat{\mathbf{S}}^2$ is diagonal in the basis of states, $|J, M_J, L, S\rangle$.
- 2 Therefore, recalling that $[\hat{S}_i, \hat{S}_j] = i\hbar\epsilon_{ijk}\hat{S}_k$ and $[\hat{L}_i, \hat{S}_k] = 0$, it follows that the matrix element of the following operator vanishes,

$$\hat{\mathbf{S}}(\hat{\mathbf{L}}\cdot\hat{\mathbf{S}}) - (\hat{\mathbf{L}}\cdot\hat{\mathbf{S}})\hat{\mathbf{S}} = \hat{L}_j[\hat{S}_i,\hat{S}_j] = i\hbar\epsilon_{ijk}\hat{L}_j\hat{S}_k \equiv -i\hbar\hat{\mathbf{S}}\times\hat{\mathbf{L}}$$

$$-i\hbar \hat{\mathbf{S}} imes \hat{\mathbf{L}} \equiv \hat{\mathbf{S}}(\hat{\mathbf{L}} \cdot \hat{\mathbf{S}}) - (\hat{\mathbf{L}} \cdot \hat{\mathbf{S}})\hat{\mathbf{S}}$$

③ Moreover, since $[\hat{\mathbf{L}} \cdot \hat{\mathbf{S}}, \hat{\mathbf{J}}] = 0$, it follows that the matrix element of the following operator also vanishes,

$$-i\hbar(\hat{\mathbf{S}}\times\hat{\mathbf{L}})\times\hat{\mathbf{J}}=\hat{\mathbf{S}}\times\hat{\mathbf{J}}(\hat{\mathbf{L}}\cdot\hat{\mathbf{S}})-(\hat{\mathbf{L}}\cdot\hat{\mathbf{S}})\hat{\mathbf{S}}\times\hat{\mathbf{J}}$$

If we expand left hand side, we thus find that the matrix element of the following operator also vanishes,

$$(\hat{\mathbf{S}} \times \hat{\mathbf{L}}) \times \hat{\mathbf{J}} = \hat{\mathbf{L}}(\hat{\mathbf{S}} \cdot \hat{\mathbf{J}}) - \hat{\mathbf{S}}(\hat{\mathbf{L}} \cdot \hat{\mathbf{J}}) \stackrel{\hat{\mathbf{L}}=\hat{\mathbf{J}}-\hat{\mathbf{S}}}{=} \hat{\mathbf{J}}(\hat{\mathbf{S}} \cdot \hat{\mathbf{J}}) - \hat{\mathbf{S}}\hat{\mathbf{J}}^2$$

5 Therefore, it follows that $\langle \hat{\mathbf{S}} \hat{\mathbf{J}}^2 \rangle = \langle \hat{\mathbf{J}} (\hat{\mathbf{S}} \cdot \hat{\mathbf{J}}) \rangle$. With $\hat{\mathbf{S}} \cdot \hat{\mathbf{J}} = \frac{1}{2} (\hat{\mathbf{J}}^2 + \hat{\mathbf{S}}^2 - \hat{\mathbf{L}}^2)$, we have $\langle \hat{S}_z \rangle \langle \hat{\mathbf{J}}^2 \rangle = \langle \hat{J}_z \rangle \langle \hat{\mathbf{S}} \cdot \hat{\mathbf{J}} \rangle$, i.e.

$$\langle \hat{S}_z \rangle = \langle \hat{J}_z \rangle \frac{J(J+1) + S(S+1) - L(L+1)}{2J(J+1)}$$

$$\langle \hat{S}_z \rangle = \langle \hat{J}_z \rangle \frac{J(J+1) + S(S+1) - L(L+1)}{2J(J+1)}$$

• As a result, we can deduce that, at first order in perturbation theory, the energy shift arising from the Zeeman term is given by

$$\Delta E_{J,M_J,L,S} = \mu_{\rm B} B \langle (\hat{J}_z + \hat{S}_z) \rangle / \hbar = \mu_{\rm B} g_J M_J B$$

where effective Landé g-factor

$$g_J = 1 + rac{J(J+1) + S(S+1) - L(L+1)}{2J(J+1)}$$

N.B. for hydrogen (S = 1/2 and $J = L \pm 1/2$), we recover previous result.

Example: atomic spectra of sodium

